

MHD for ASTROPHYSICISTS

→
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NOTES: <http://www.astro.iag.usp.br/~dalpino/?q=teaching>
(references therein)

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Plasmas in Astrophysics

Class 1

Contents

- What is plasma?
- Why plasmas in astrophysics?
- Quasi-neutrality

What is PLASMA?

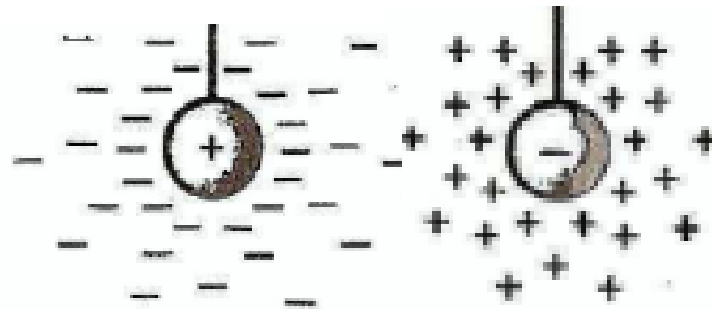
- Gas with sufficient number of **free charged particles** (positive + negative) → so that its behaviour dominated by **electromagnetic forces**.
- Even **low ionization** degree ($n_{\text{ch}}/n_{\text{total}} \sim 1\%$) : sufficient for gas to show electromagnetic properties (electrical conductivity $\sim$ fully ionized gas) -> **PLASMA!**

More formal PLASMA Concept

Plasma: quasi-neutral gas of charged particles (electrons+ions):

$$e n_e \sim e Z n_i$$

with long range *collective interactions*:



(binary collisions not important)

Why Plasmas in Astrophysics ?

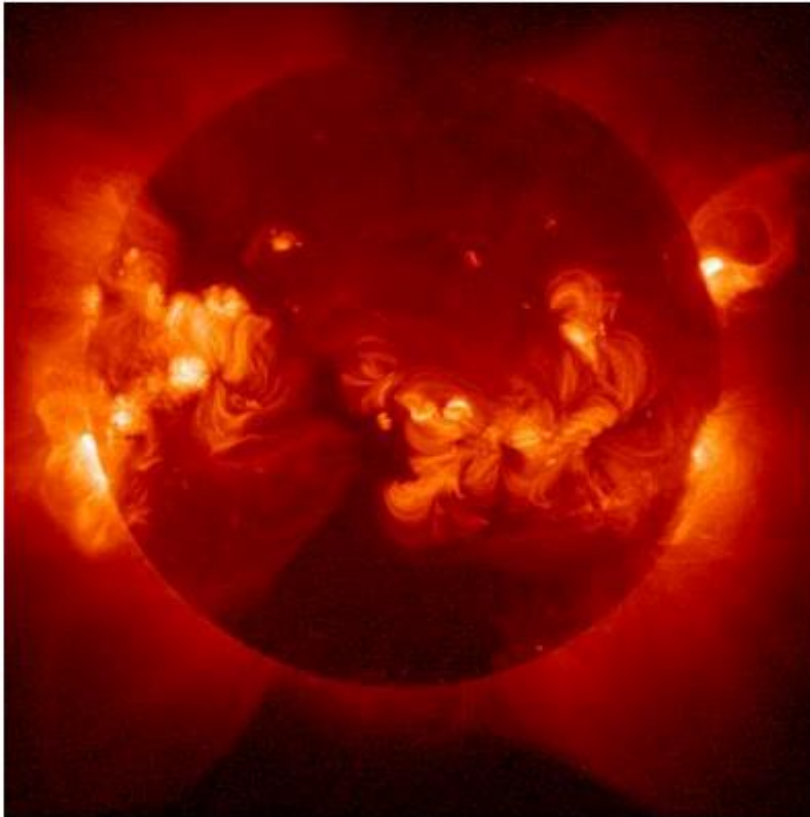
The universe does not consist of ordinary matter

Over 90% of visible matter in Universe:

PLASMA !

Exs. Astrophysical Plasmas

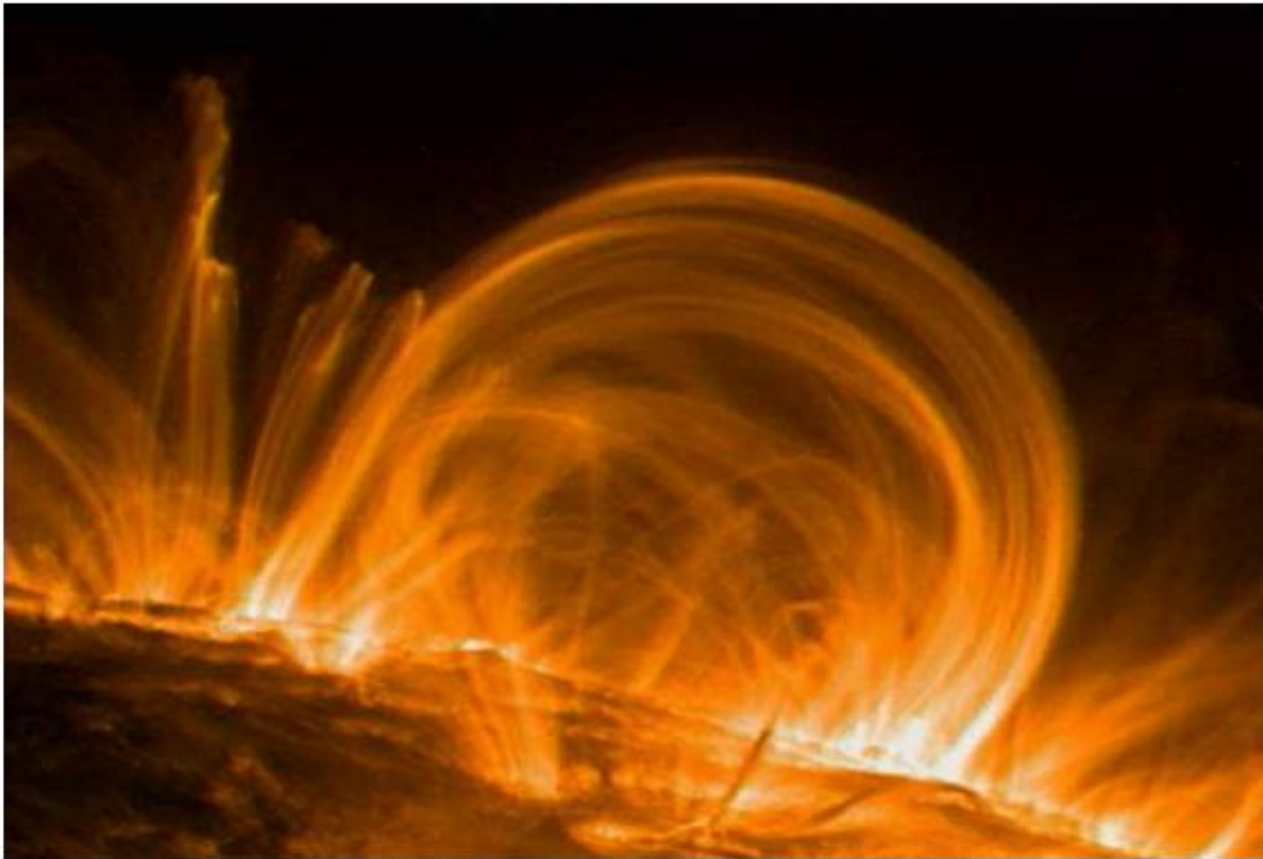
Example: The Sun



a magnetized plasma!
(sunatallwavelengths.mpeg)

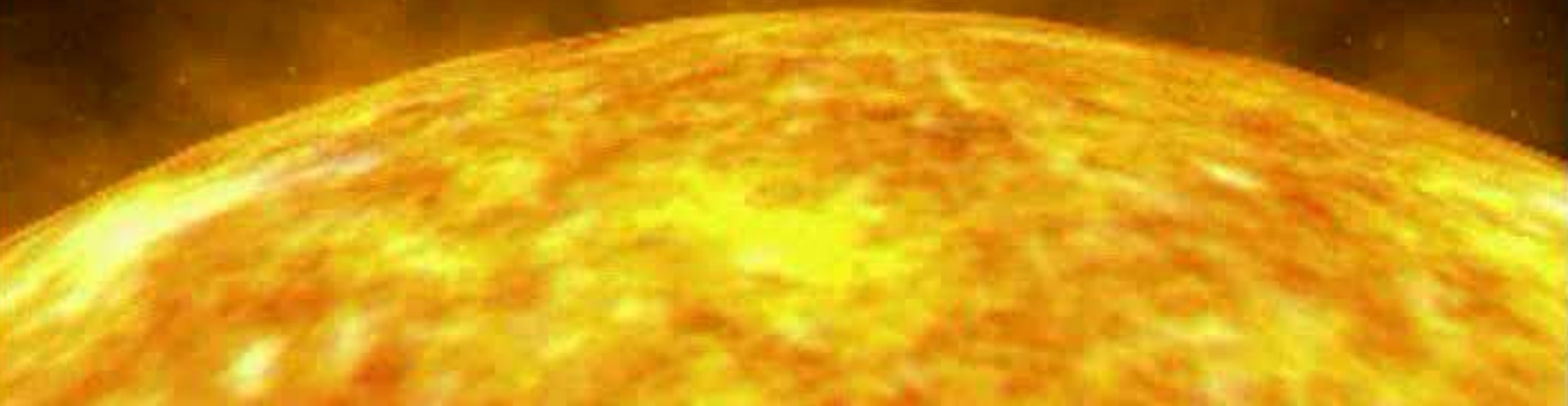
Exs. Astrophysical Plasmas

Example: Coronal loops (cont'd)

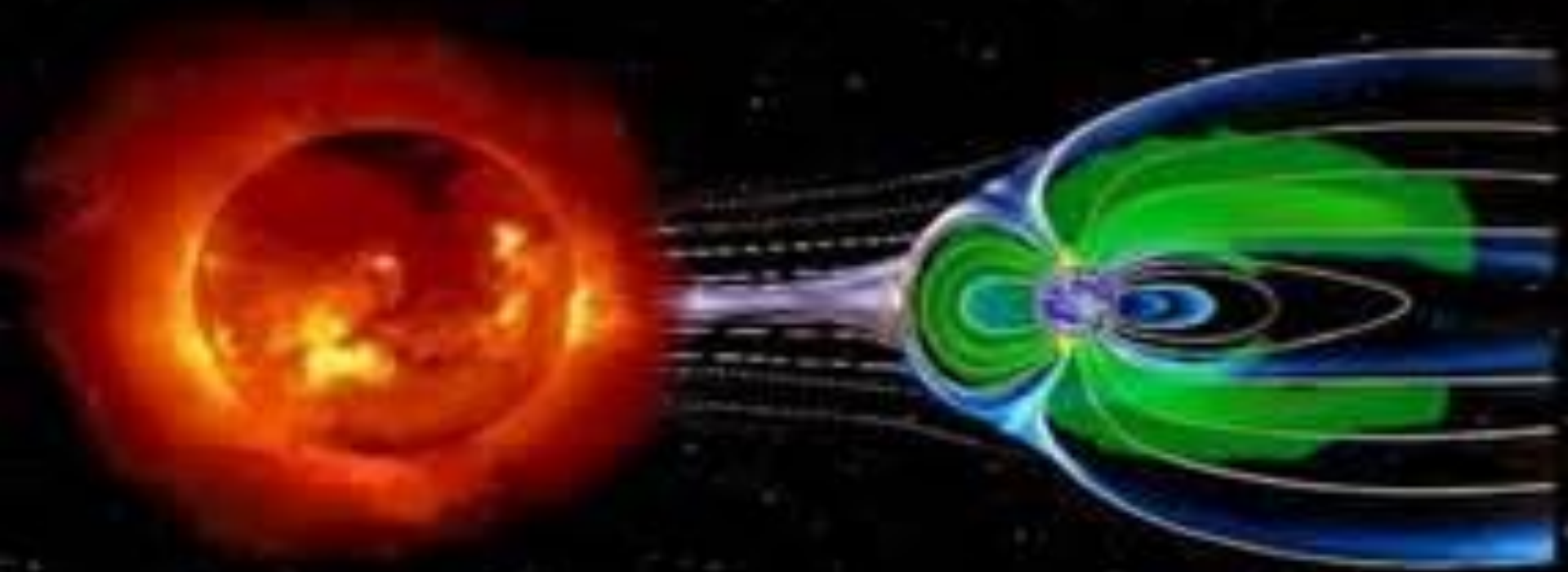


[from recent observations with TRACE spacecraft]

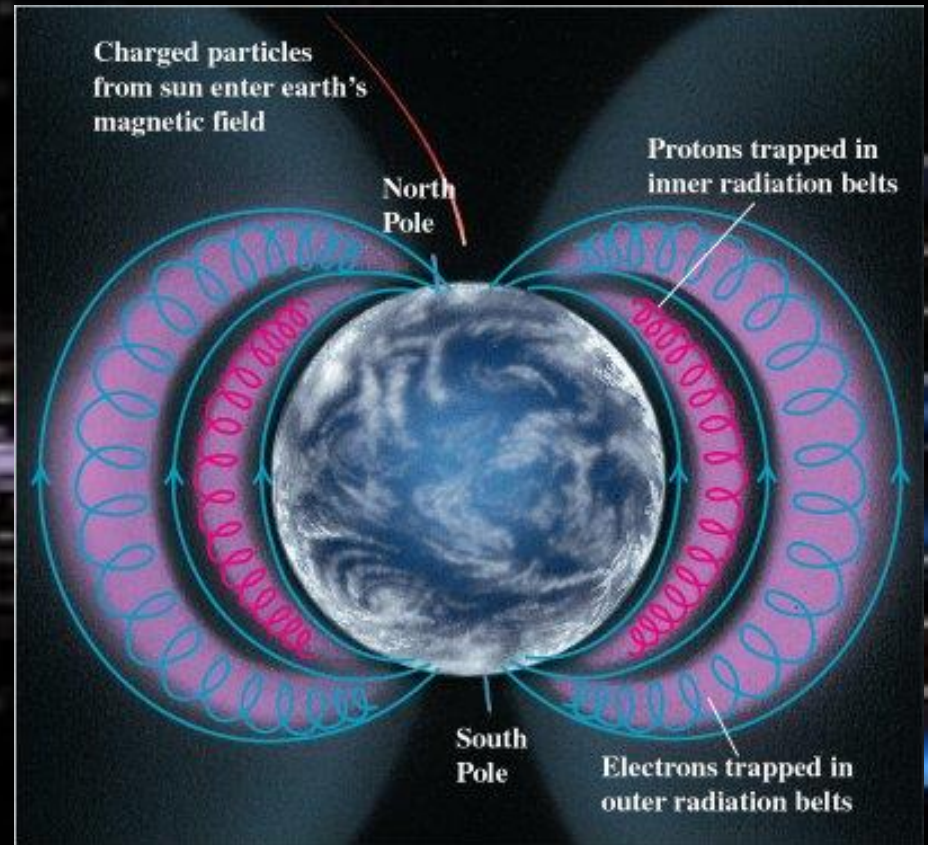
Sun and Earth



Ex. Plasmas in Earth

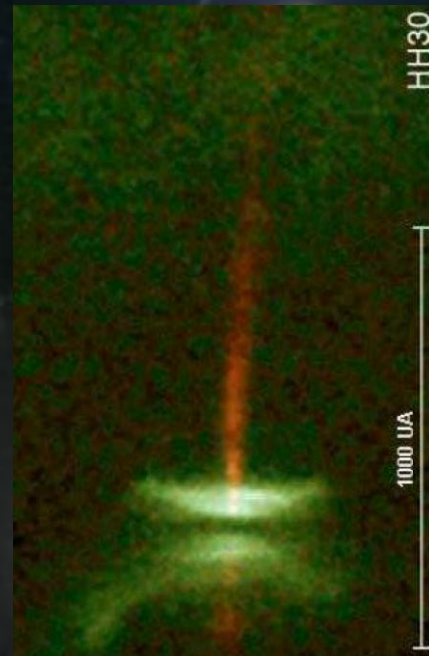


Ex. Plasmas in Earth



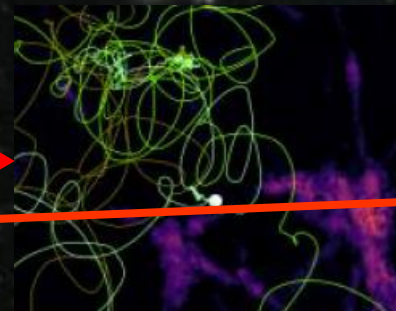
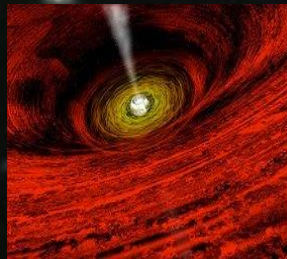
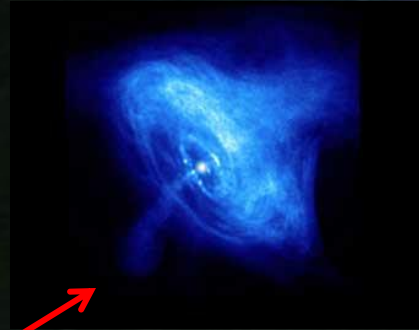
Importance of Plasmas

- Magnetized plasmas are present in almost all astrophysical objects
- They are **crucial** in:
 - star formation; late stages
 - solar and stellar activity
 - formation of jets and accretion disks
 - formation and propagation of cosmic rays
 - galaxy structure



Importance of Plasmas

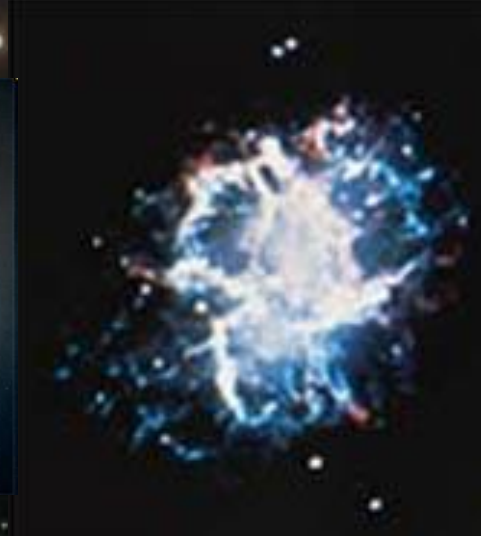
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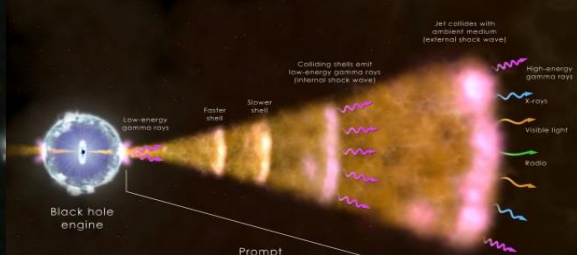
Importance of Plasmas

- They are also **crucial** in:

- ISM
- molecular clouds
- supernova remnants
- proto-planetary disks
- planetary nebulae



- GRBs



Importance of Plasmas

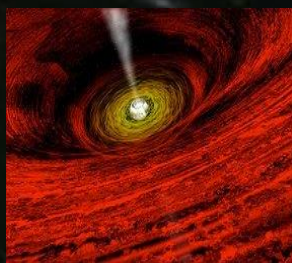
- Their importance ***not well understood*** yet in:

- stellar evolution

- galaxy evolution

- structure formation in the early Universe





Black Holes surrounds



Accretion
disks

Sun & Stars



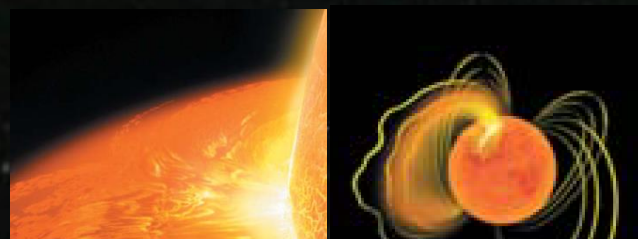
Star Formation
and ISM



Pulsars

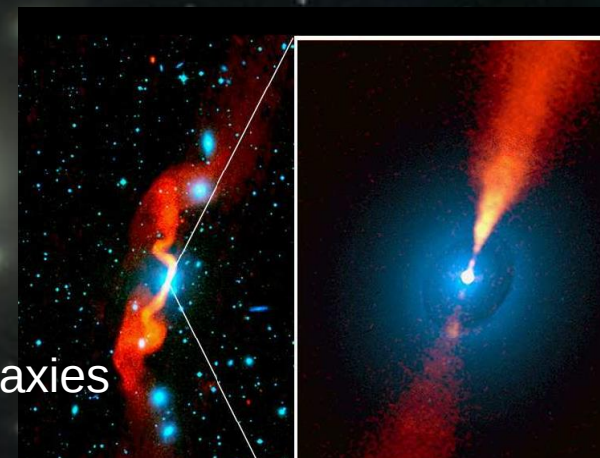


Neutron Stars

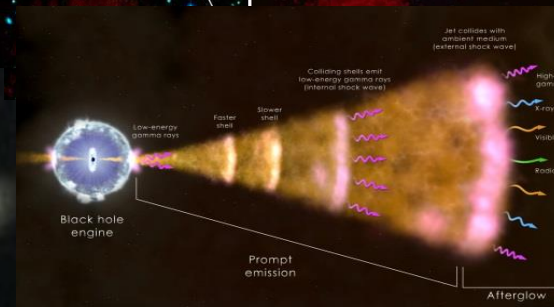


**PLASMAS
in
Astrophysics**

Galaxies



Gamma-Ray Bursts

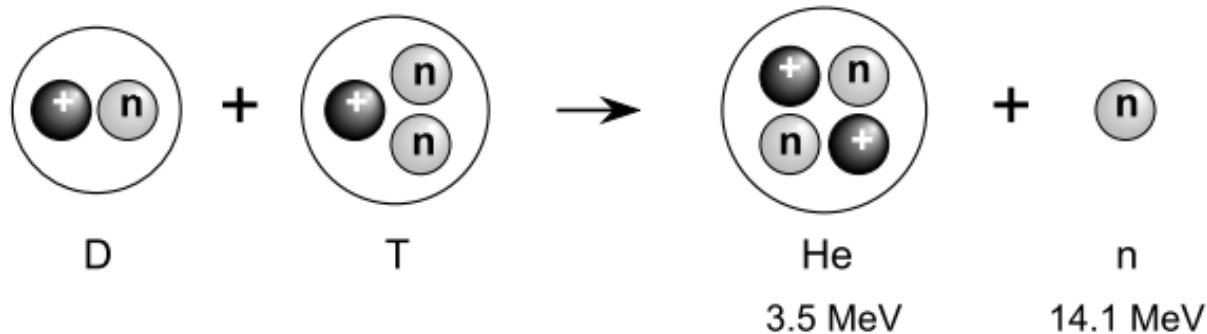


Plasmas also in Laboratory

Major goal is to produce clean energy via:

- Thermonuclear fusion (**as in the interior of stars**) → production of energy!

Reactions of hydrogen isotopes



Magnetic Fields in a PLASMA provide the only way to confine matter of very high temperatures during long times!

Understanding Quasi Neutrality

$$Zen_i = en_e$$

- Let us assume a gas of charged particles with local charge concentration ($e \delta n_e$)

→ according to Coulomb law: $\vec{\nabla} \cdot \vec{E} = \frac{\partial E_{//}}{\partial x} = -4\pi e \delta n_e$

→ generates the electric field: $E_{//} = 4\pi e n_e \xi$

→ $E_{//}$ provokes on the thermal random motion of electrons: flow with velocity $v_{//}$:

$$\frac{dv_{//}}{dt} = \frac{d^2 \xi}{dt^2} = -\frac{e}{m_e} E_{//} = -\frac{4\pi n_e e^2}{m_e} \xi$$

- Solution: simple harmonic motion with plasma electron frequency that in the average neutralizes $E_{//}$:

$$\omega_{pe} = \left(\frac{4\pi n_e e^2}{m_e} \right)^{1/2} \simeq 5.6 \times 10^4 n_e^{1/2} \text{ s}^{-1}$$

Quasi Neutrality- Plasma Frequency

- ω_{pe} defines natural plasma frequency (neutralizes E effect)

- ions oscillate with much $\ll$ frequency $\rightarrow$
(practically at rest in relation to els.)

$$\omega_{pi} = \left(\frac{Zm_e}{m_i} \right)^{1/2} \omega_{pe}$$

- Any external electric field E applied with freq. $\omega < \omega_{pe}$ is unable to penetrate the plasma

- Debye length: $\lambda_D = v_{th}/\omega_{pe}$

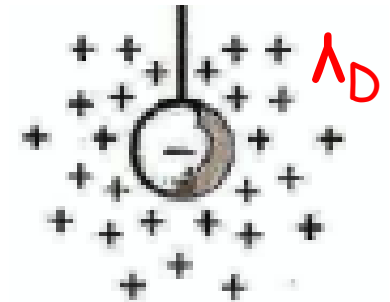
$\rightarrow E$ does not penetrate plasma if:

$$\omega < \omega_{pe} \rightarrow \lambda > \lambda_D$$

PLASMA: Debye length (λ_D)

- $\lambda_D \rightarrow$ scale within which separation between charges can be "felt":

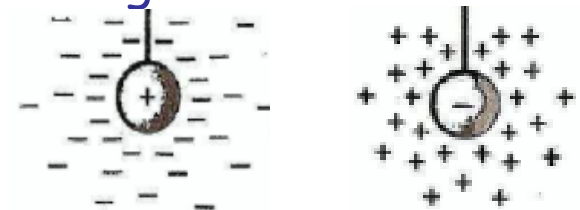
$$\lambda_D \simeq \frac{v_{th}}{\omega_{pe}} \simeq \left(\frac{KT_e}{4\pi n_e e^2} \right)^{1/2} = 7 \text{ cm } (T_e/n_e)^{1/2}$$



- Within sphere of radius (λ_D) $\rightarrow$ charge neutrality is not valid: electrostatic external oscillations with $\lambda \lesssim \lambda_D$ penetrate the sphere and feel the collective effects of the charges $\rightarrow$ strongly damped (Landau damping)

- Electric Potential Field of a charge within plasma: has its action screened (or partially blocked) by clouds of charges

$$\phi(r) = \frac{-en_1}{r} e^{-r/\lambda_D}$$



Plasma Quasi Neutrality

$$Zen_i = en_e$$

IS VALID WHEN:



$$\omega < \omega_{pe} \text{ or } \lambda > \lambda_D$$

Quasi Neutrality in Astrophysical Plasmas

- Typical dimension of astrophysical plasmas: $L \gg \lambda_D$

⇒ quase neutrality is valid

→ Internal E fields: little important !

(neutralized by strong plasma oscillations ω_{pe})

→ External E fields typically do not penetrate

$$(\omega \ll \omega_{pe}, \lambda \gg \lambda_D)$$

Plasma: collective behaviour

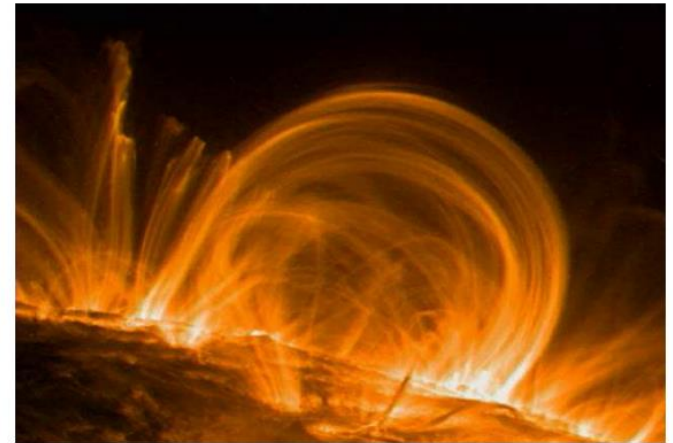
$$\lambda_D \simeq \frac{v_{th}}{\omega_{pe}} \simeq \left(\frac{KT_e}{4\pi n_e e^2} \right)^{1/2} = 7 \text{ cm } (T_e/n_e)^{1/2}$$

Ex. Solar corona:



$$T_e \sim 10^6 \text{ K}, n_e \sim 10^{10} \text{ cm}^{-3}$$
$$\rightarrow \lambda \gg \lambda_D \sim 0.07 \text{ cm}$$

Example: Coronal loops (cont'd)



[from recent observations with TRACE spacecraft]

-> So that the system size λ is much larger than λ_D : collective behaviour

Plasma Physics

Three theoretical models:

- **Theory of motion of single charged particles** in given magnetic and electric fields;
- **Kinetic theory of a collection of such particles**, describing plasmas *microscopically* by means of particle distribution functions $f_{e,i}(\mathbf{r}, \mathbf{v}, t)$;
- **Fluid theory (magnetohydrodynamics)**, describing plasmas in terms of averaged *macroscopic* functions of $\mathbf{r}$ and t .

Plasma Physics

Three theoretical models:



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Motion of a single charge

Equation of motion

of charged particle in given electric and magnetic field, $\mathbf{E}(\mathbf{r}, t)$ and $\mathbf{B}(\mathbf{r}, t)$:

$$m \frac{d\vec{v}}{dt} = q \left(\vec{E} + \frac{\vec{v}}{c} \times \vec{B} \right)$$

Let us

- Apply to constant magnetic field $\mathbf{B} = B\mathbf{e}_z$, $\mathbf{E} = 0$:

- Systematic solution of Eq. (1) with $\mathbf{v} = d\mathbf{r}/dt = (\dot{x}, \dot{y}, \dot{z})$ gives two coupled differential equations for motion in the perpendicular plane:

$$\ddot{x} - (qB/mc) \dot{y} = 0,$$

$$\ddot{y} + (qB/mc) \dot{x} = 0.$$

$\Rightarrow$ periodic motion about a fixed point $x = x_c, y = y_c$ (the guiding centre).

CGS

Motion of a single charge



Cyclotron motion

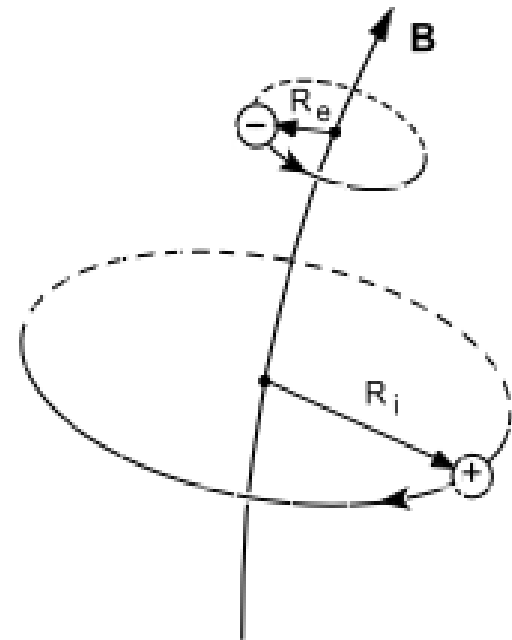
This yields periodic motion in a magnetic field, with *gyro- (cyclotron) frequency*

$$\omega_B = \frac{|q| B}{mc} = \Omega$$

and *cyclotron (gyro-)radius*

$$r_B = \frac{v_{\perp}}{\omega_B} \approx \frac{\sqrt{2mkT}}{|q|B} c$$

⇒ Effectively, charged particles stick to the field lines.



Opposite motion of electrons and ions about guiding centres with quite different gyro-frequencies and radii, since $m_e \ll m_i$:

$$\Omega_e \equiv \frac{eB}{m_e c} \gg \Omega_i \equiv \frac{ZeB}{m_i c}, \quad R_e \approx \frac{\sqrt{2m_e kT}}{eB} c \ll R_i \approx \frac{\sqrt{2m_i kT}}{ZeB} c$$

Motion of a single charge

Cyclotron motion (cont'd)

Orders of magnitude

- Typical gyro-frequencies, e.g. for **corona** plasma $B \sim 10^4$ G:

$$\Omega_e = 5.3 \times 10^{11} \text{ rad s}^{-1}$$

$$\Omega_i = 2.9 \times 10^8 \text{ rad s}^{-1}$$

- Gyro-radii, with $v_{\perp} = v_{\text{th}} \equiv \sqrt{2kT/m}$ for $T_e = T_i \sim 10^6$ K:

$$v_{\text{th},e} \sim 3 \cdot 10^8 \text{ cm/s} \qquad R_e \sim 10^{-3} \text{ cm}$$

$$v_{\text{th},i} \sim 10^7 \text{ cm/s} \qquad \Rightarrow \qquad R_i \sim 3 \cdot 10^{-2} \text{ cm}$$

$\Rightarrow$ The large corona time scales and dimensions *justify averaging*.

Motion of a single charge

- Magnetic fields:
 1. charged particles gyrate around field lines;
 2. fluid and magnetic field move together (" B frozen into the plasma");

Plasma Physics

Three theoretical models:

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Kinetic Description of a Plasma

A plasma consists of a very large number of interacting charged particles $\Rightarrow$ *kinetic plasma theory* derives the equations describing the *collective behavior* of the many charged particles by applying the methods of statistical mechanics.



$$(\omega \geq \omega_{pe} \quad l \geq \lambda_D)$$

Kinetic theory describes the collective behaviour of the many charged particles by means of particle distribution functions (**Fokker-Planck equation**):

$f(r, u, t)$ $\rightarrow$

$$\frac{\partial f}{\partial t} + \mathbf{u} \cdot \nabla_x f + \frac{q}{m} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_u f = 0$$

Plasma Physics

Three theoretical models:

- **Theory of motion of single charged particles** in given magnetic and electric fields;
- **Kinetic theory of a collection of such particles**, describing plasmas *microscopically* by means of particle distribution functions $f_{e,i}(\mathbf{r}, \mathbf{v}, t)$;



Fluid theory (magnetohydrodynamics), describing plasmas in terms of averaged *macroscopic* functions of $\mathbf{r}$ and t .

Fluid description: Magnetohydrodynamics (MHD)

Macroscopic definition

Macroscopic model: size and time scales are large enough $\rightarrow$ possible to apply AVERAGES over microscopic quantities: collective plasma oscillations and collective cyclotron motions of ions and electrons

(a) $\tau \gg \Omega_i^{-1} \sim B^{-1}$ (time scale longer than inverse cyclotron frequency);

(b) $\lambda \gg R_i \sim B^{-1}$ (length scale larger than cyclotron radius).

$\Rightarrow$ MHD $\equiv$ magnetohydrodynamics

PLASMA @ different scales

Intermediate - *microscopic scales* : $l \gg \lambda_D$

kinetic theory describes the collective behaviour of the many charged particles by means of particle distribution functions (**Fokker-Planck equation**):

$$f_{e,i}(r, v, t)$$

Large – *macroscopic scales* : $L \gg R_L$

effectively **collisional** -> **fluid description**: size and time scales are large enough → possible to apply AVERAGES over microscopic quantities: over collective plasma oscillations and collective cyclotron motions - **MHD equations**



Applicable to most astrophysical plasmas

End of Class 1